

**Erratum: “Resolved Nuclear Kinematics Link the Formation and Growth of Nuclear Star Clusters with the Evolution of Their Early and Late-type Hosts” (2021, ApJ, 921, 8)**

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Received 2021 December 2; published 2021 December 29

We have identified an error in Equations (D3) and (D4) in Appendix D of the published article. These equations result from the combination of Equations (16), (17), and (18) in Harborne et al. (2020). In the published article, the right-hand sides are swapped with each other. The corrected versions of the equations are

$$\Delta\left(\frac{V}{\sigma}\right)_e^{\text{corr}} = \left\{ \frac{7.55}{1 + \exp\left[4.42\left(\frac{\sigma_{\text{PSF}}}{R_e^{\text{maj}}}\right)^{1.55} + 2.73\right]} - 0.46 \right\} \quad (\text{D3})$$

$$+ 3\left(\frac{\sigma_{\text{PSF}}}{R_e^{\text{maj}}}\right) \times [-0.10\epsilon_e + 0.024\log_{10}(n_{\text{NSC}}) + 0.064]$$

and

$$\Delta\lambda_{R_e}^{\text{corr}} = \left\{ \frac{7.48}{1 + \exp\left[4.08\left(\frac{\sigma_{\text{PSF}}}{R_e^{\text{maj}}}\right)^{1.60} + 2.89\right]} - 0.39 \right\} \quad (\text{D4})$$

$$+ \left(\frac{\sigma_{\text{PSF}}}{R_e^{\text{maj}}}\right) \times [0.10\epsilon_e - 0.22\log_{10}(n_{\text{NSC}}) + 0.10].$$

All results shown and discussed in the published article were obtained using the correct version of the equations and do not need any modifications.

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References

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